

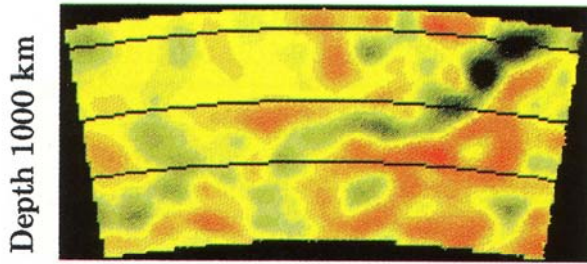
Playing dice with the Earth's interior

Jeannot Trampert
Utrecht University

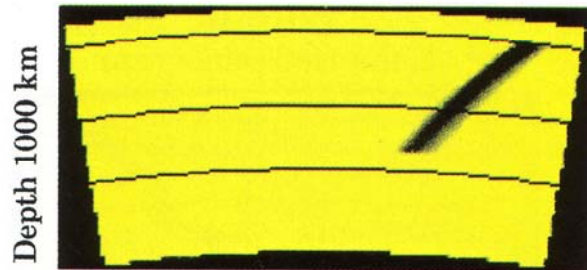


Thanks to many excellent PhD students and postdocs

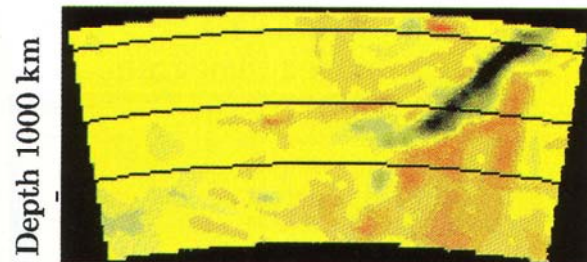
Initial tomography



Theoretical slab



Enhanced tomogram



Surface distance 2200 km



Is this the only model compatible with the data?

Project some chosen model onto the model null space of the forward operator

Then add to original model.
The same data fit is guaranteed!

True model

Forward problem
(exact theory)

Data

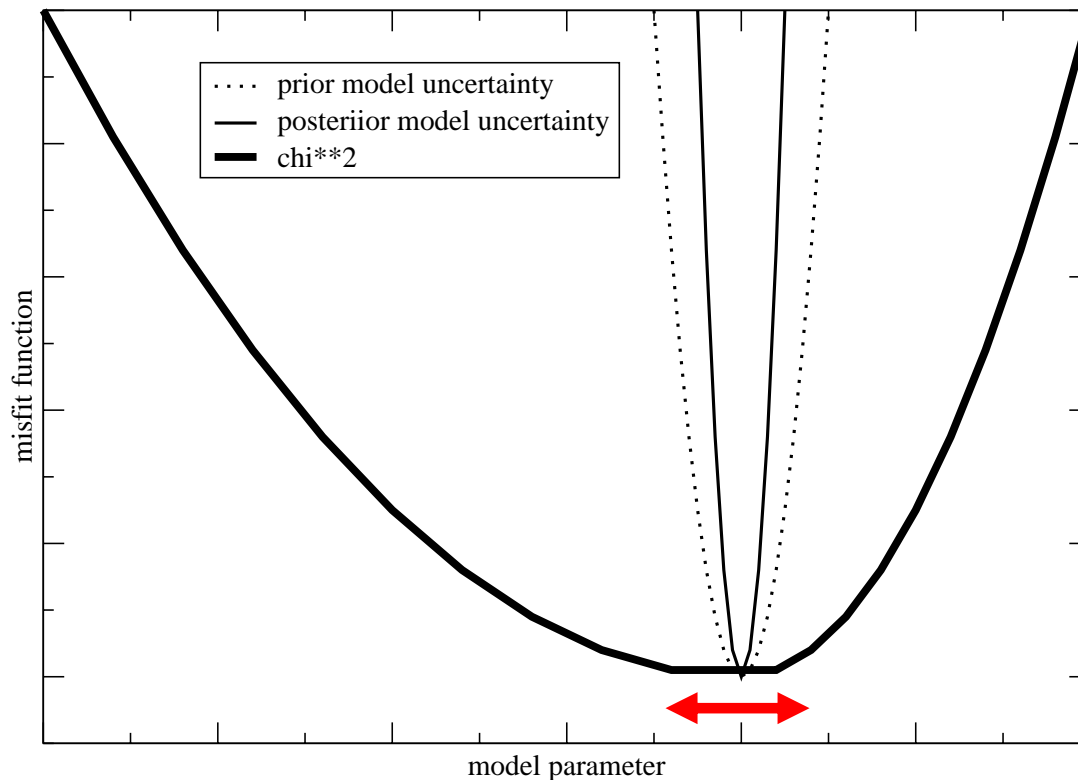
APPRAISAL
Model uncertainty

Estimated model

Inverse problem
(approximate theory, uneven data
coverage, data errors)

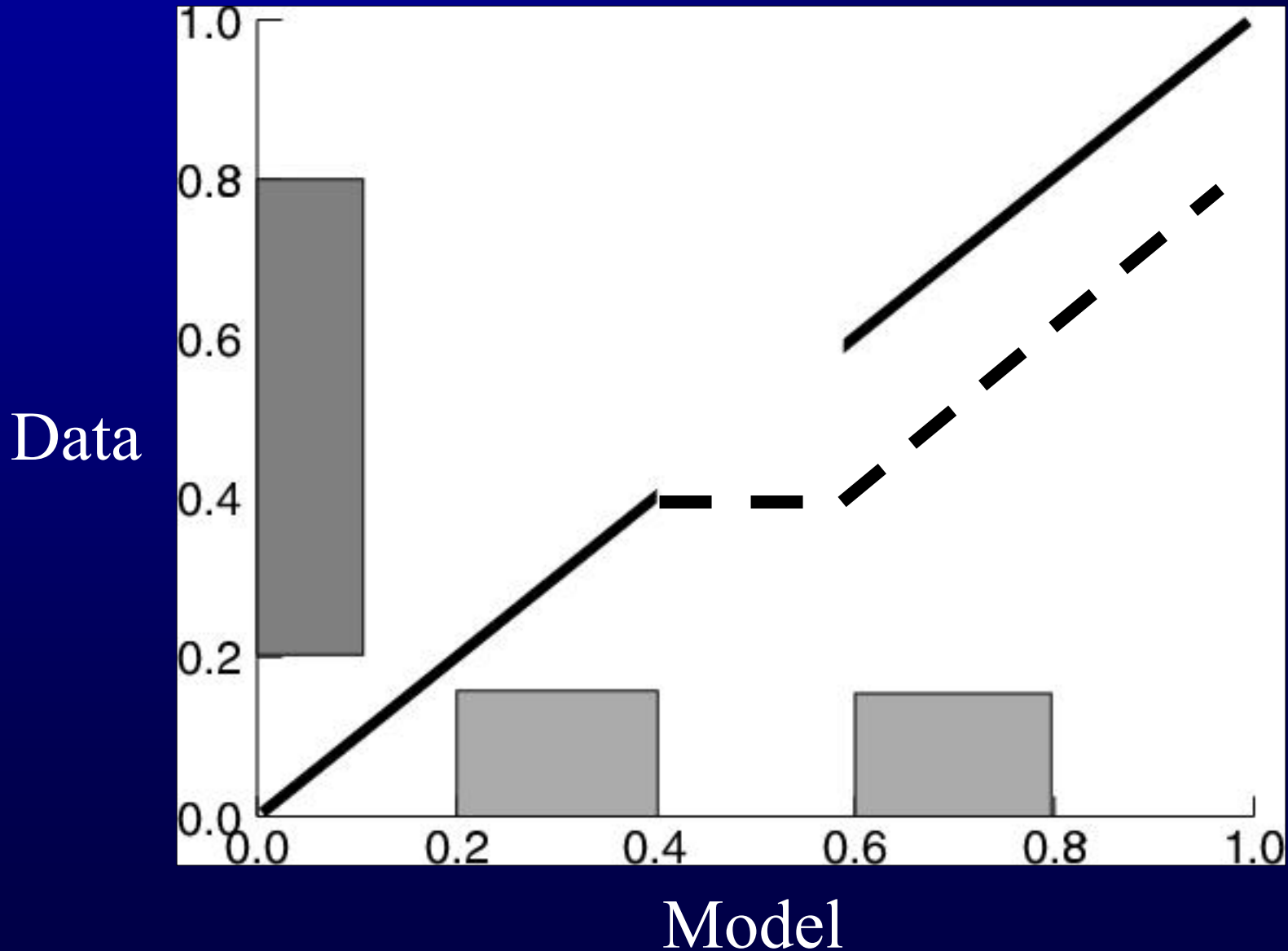
- Ill-posed (null space, continuity)
- Ill-conditioned (error propagation)

In the presence of a model null space, the cost function has a flat valley and the standard least-squares analysis underestimates uncertainty.



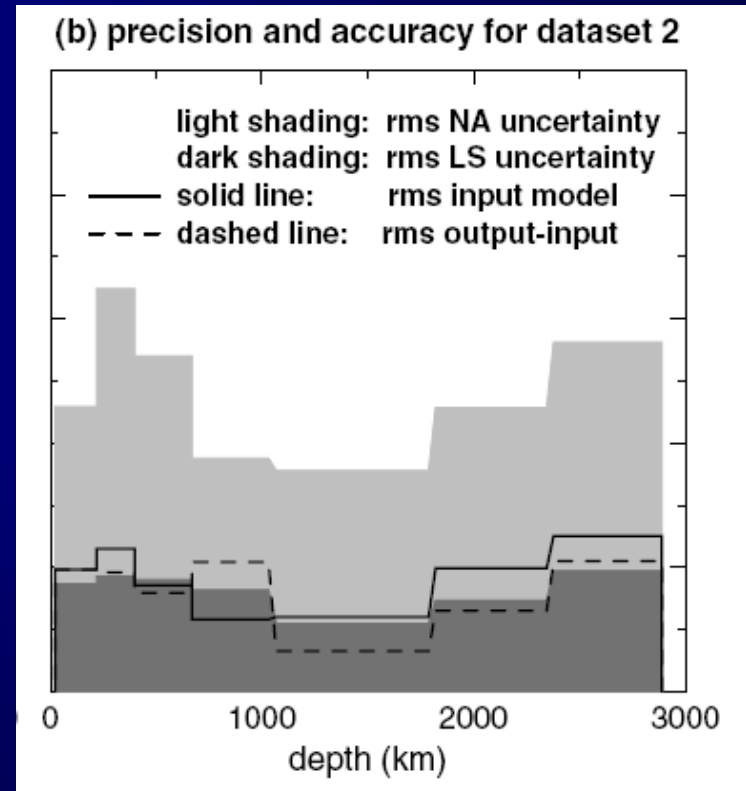
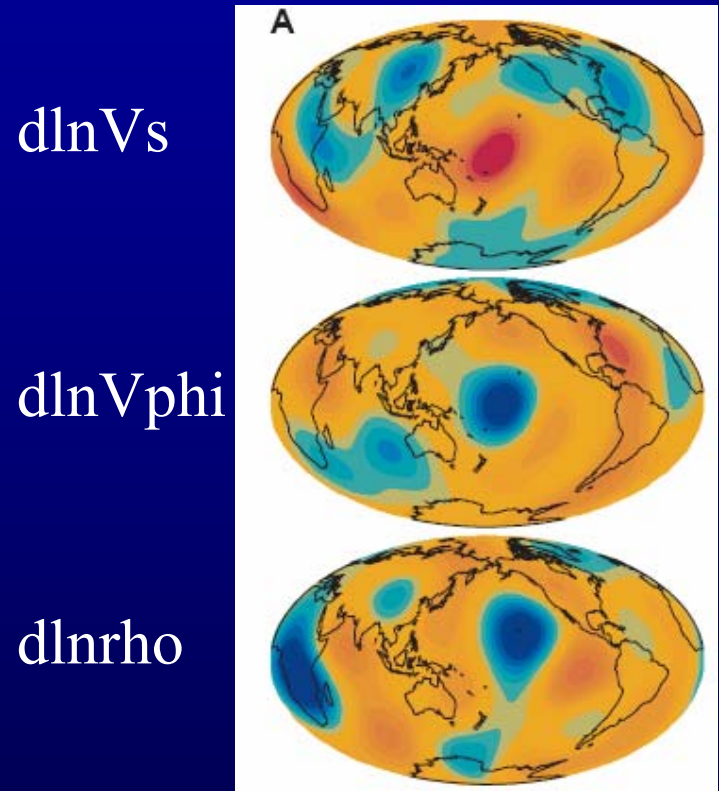
Realistic uncertainty for the given data

Multiple solutions are possible if the model does not continuously depend on the data, even for a linear problem.



Does it matter?

Yes!



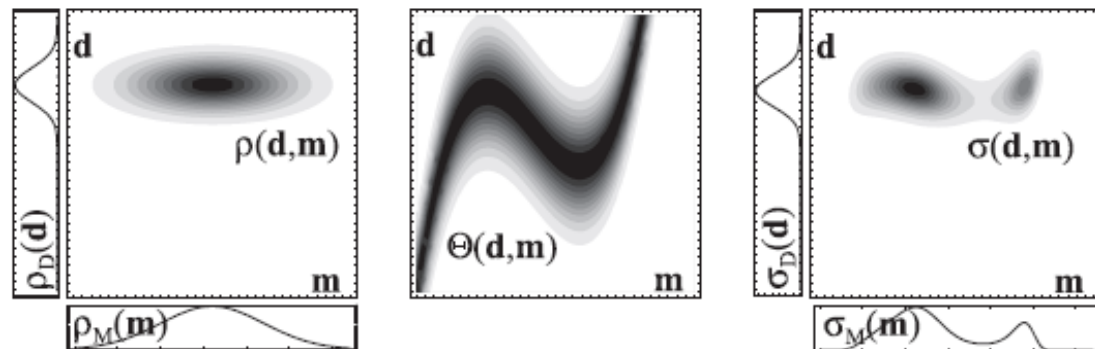
SPRD6 (Ishii and Tromp, 1999)

Input model is density from SPRD6 (Resovsky and Trampert, 2002)

The most general solution of an inverse problem (Bayes, Tarantola)

$$\sigma(d, m) = k \frac{\rho(d, m) \Theta(d, m)}{\mu(d, m)}$$

$$\text{posterior} = \frac{\text{prior} \times \text{likelihood}}{\text{evidence}}$$



A full model space search should estimate

$$\sigma(m) = k\rho_m(m)L(m)$$

1. Exhaustive search
2. Brute force Monte Carlo (Shapiro and Ritzwoller, 2002)
3. Simulated Annealing (global optimisation with convergence proof)
4. Genetic algorithms (global optimisation with no convergence proof)
5. Sample $\rho(m)$ and apply Metropolis rule on $L(m)$. This will result in importance sampling of $\sigma(m)$ (Mosegaard and Tarantola, 1995)
6. Neighbourhood algorithm (Sambridge, 1999)
7. Neural networks (Meier et al., 2007)

The model space is HUGE!

Draw a 1000 models per second where
 $m = \{0, 1\}$

$M=30 \rightarrow 13$ days

$M=50 \rightarrow 35702$ years

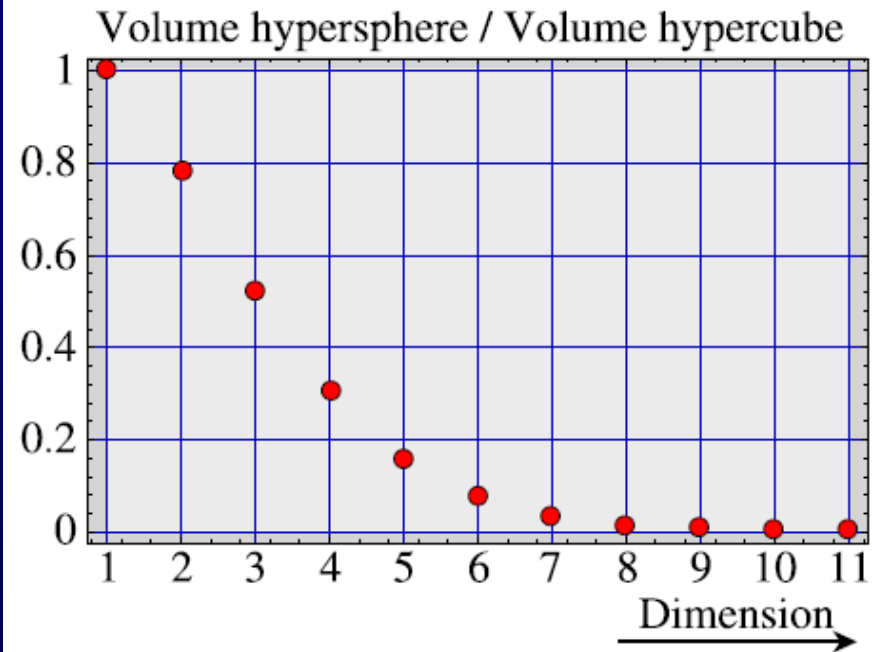
Seismic tomography $M = O(1000-100000)$

The model space is EMPTY!



$$\text{Volume hypersphere: } \frac{2 \pi^{n/2} r^n}{n \Gamma(n/2)}$$

$$\text{Volume hypercube: } (2r)^n$$



Tarantola, 2005

The curse of dimensionality

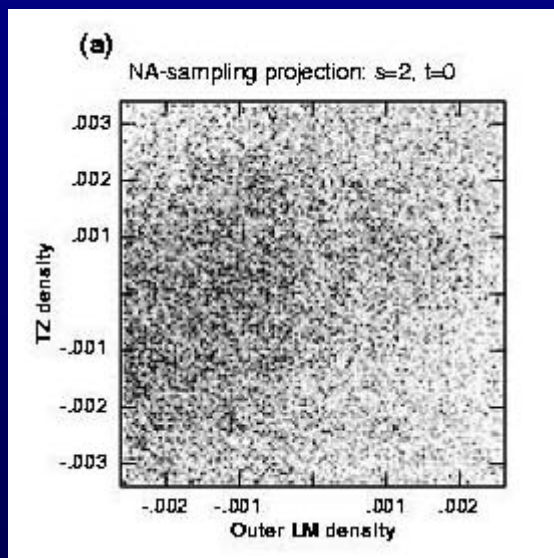
→ small problems $M \sim 30$

$$\sigma(m) = k \rho_m(m) L(m)$$

1. Exhaustive search, **IMPOSSIBLE**
2. Brute force Monte Carlo, **DIFFICULT**
3. Simulated Annealing (global optimisation with convergence proof), **DIFFICULT**
4. Genetic algorithms (global optimisation with no convergence proof), ???
5. Sample $\rho(m)$ and apply Metropolis rule on $L(m)$. This will result in importance sampling of $\sigma(m)$ (Mosegaard and Tarantola, 1995), **BETTER, BUT HOW TO GET A MARGINAL**
6. Neighbourhood algorithm (Sambridge, 1999), **THIS WORKS**
7. Neural networks, **PROMISING**

The neighbourhood algorithm:

Sambridge 1999



Stage 1:

Guided sampling of the model space.

Samples concentrate in areas (neighbourhoods) of better fit.

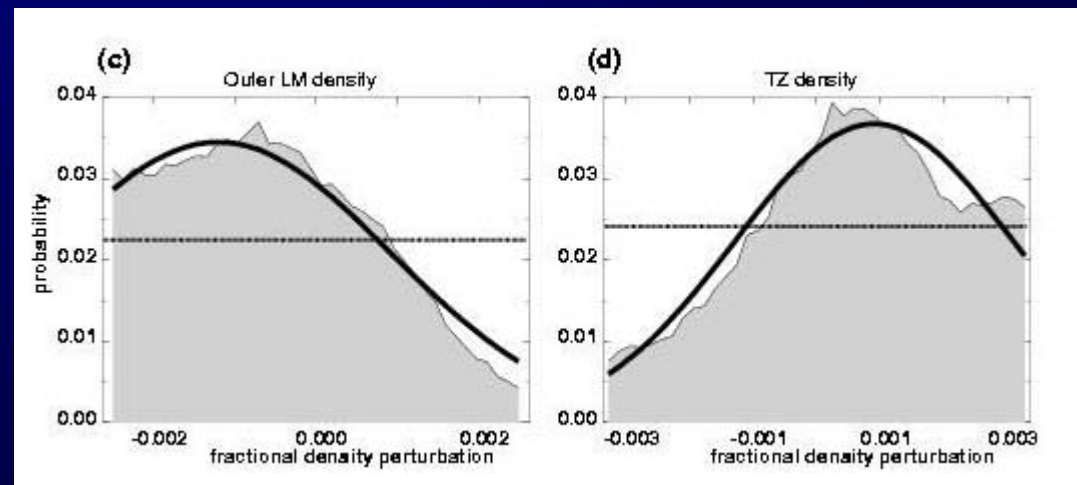
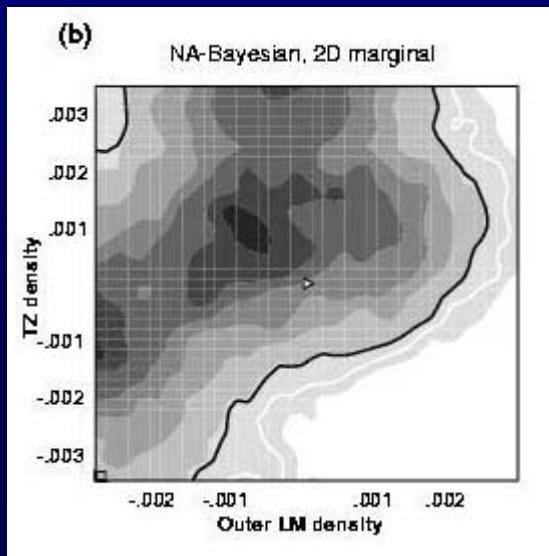
The neighbourhood algorithm (NA):

Stage 2: importance sampling

Resampling so that sampling density reflects posterior

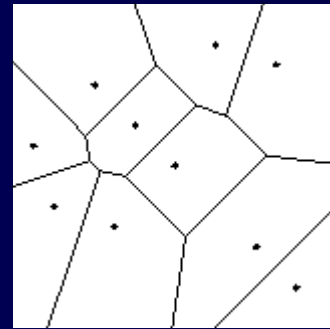
2D marginal

1D marginal

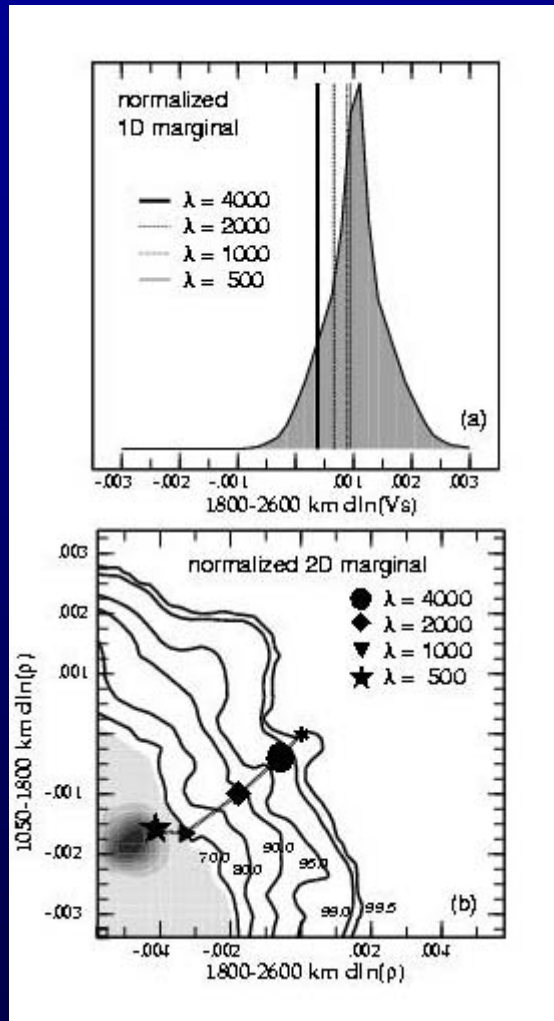


Advantages of NA

1. Interpolation in model space with Voronoi cells
2. Relative ranking in both stages (less dependent on data uncertainty)
3. Marginals calculated by Monte Carlo integration \rightarrow convergence check



NA and Least Squares consistency



As damping is reduced, LS solution converges towards most likely NA solution.

In the presence of a null space, LS solution will diverge, but NA solution remains unaffected.

Finally some tomography!

We sampled all models compatible with the data using the **Neighbourhood Algorithm** (Sambridge, GJI 1999)

Data: **649 modes** (NM, SW fundamentals and overtones)

He and Tromp, Laske and Masters, Masters and Laske, Resovsky and Ritzwoller, Resovsky and Pestena, Trampert and Woodhouse, Tromp and Zanzerkia, Woodhouse and Wong, van Heijst and Woodhouse, Widmer-Schmidrig

Parameterization: 5 layers

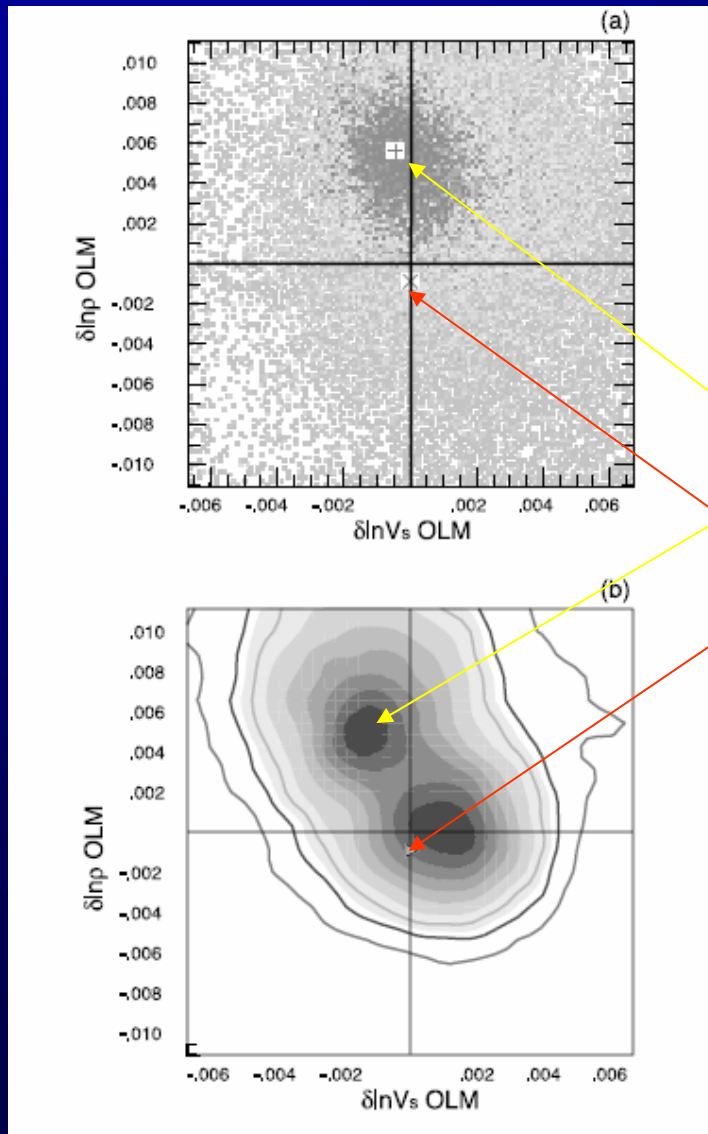
[**2891-2000 km**] [**2000-1200 km**] [**1200-660 km**]

[660-400 km] [400-24 km]

Seismic parameters

$d\ln V_s$, $d\ln V_\Phi$, $d\ln \rho$, relative topography on CMB and 660

Back to density



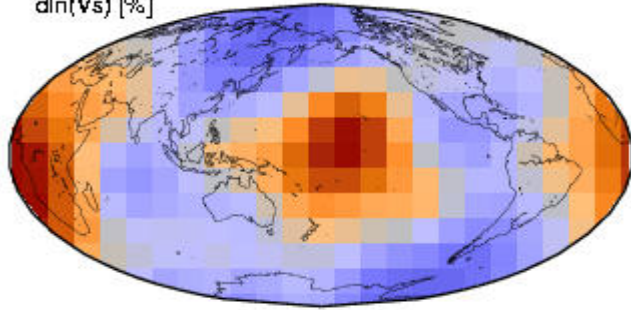
Stage 1

Most likely model
SPRD6

Stage 2

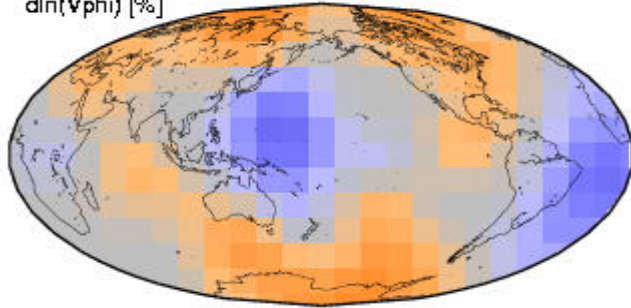
2000-2981 km

$d\ln(V_s)$ [%]



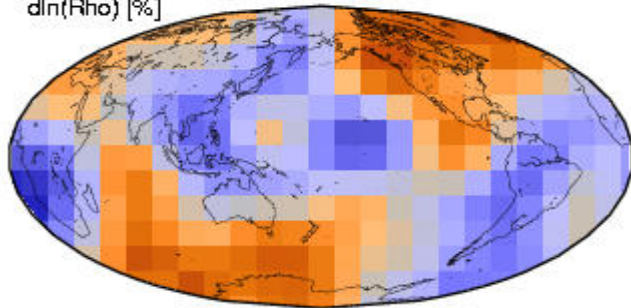
-1 0

$d\ln(V_{phi})$ [%]



-1 0

$d\ln(\rho)$ [%]



-1 0

➤ Gravity filtered models over 15×15 degree equal area caps

➤ Likelihoods in each cap are nearly Gaussian

➤ Most likely model (above one standard deviation)

➤ Uniform uncertainties

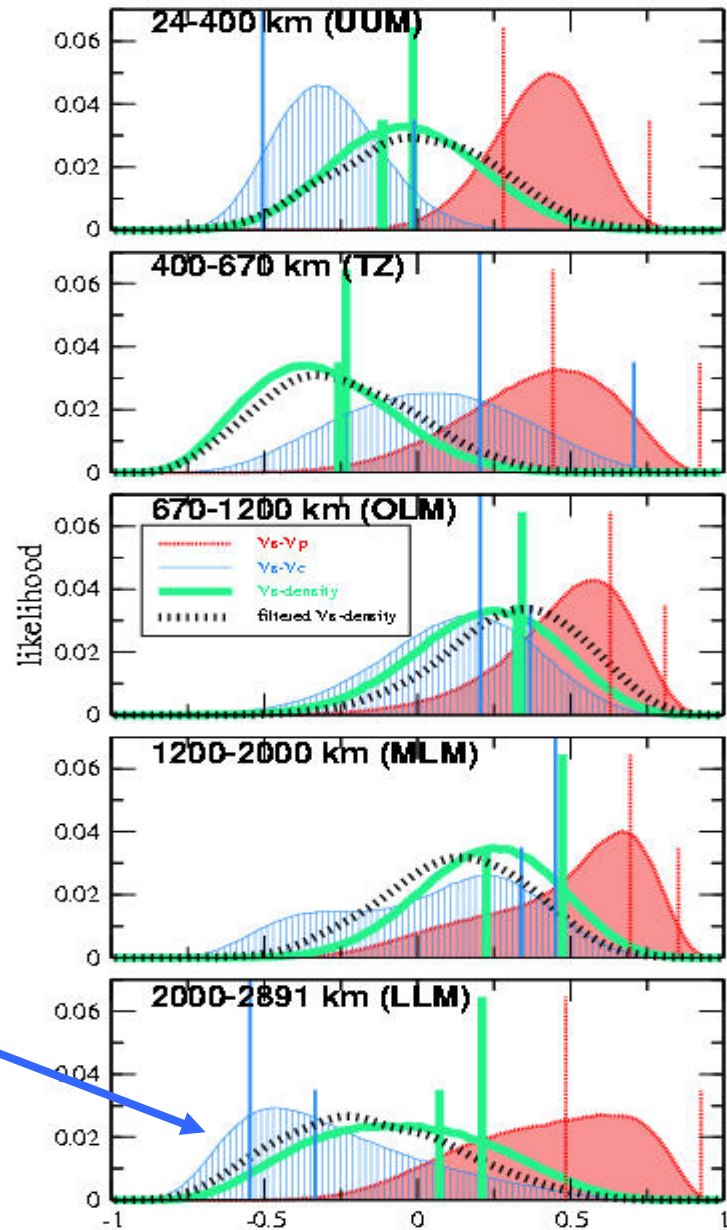
▪ $d\ln V_s = 0.12$ %

▪ $d\ln V_{\Phi} = 0.26$ %

▪ $d\ln \rho = 0.48$ %

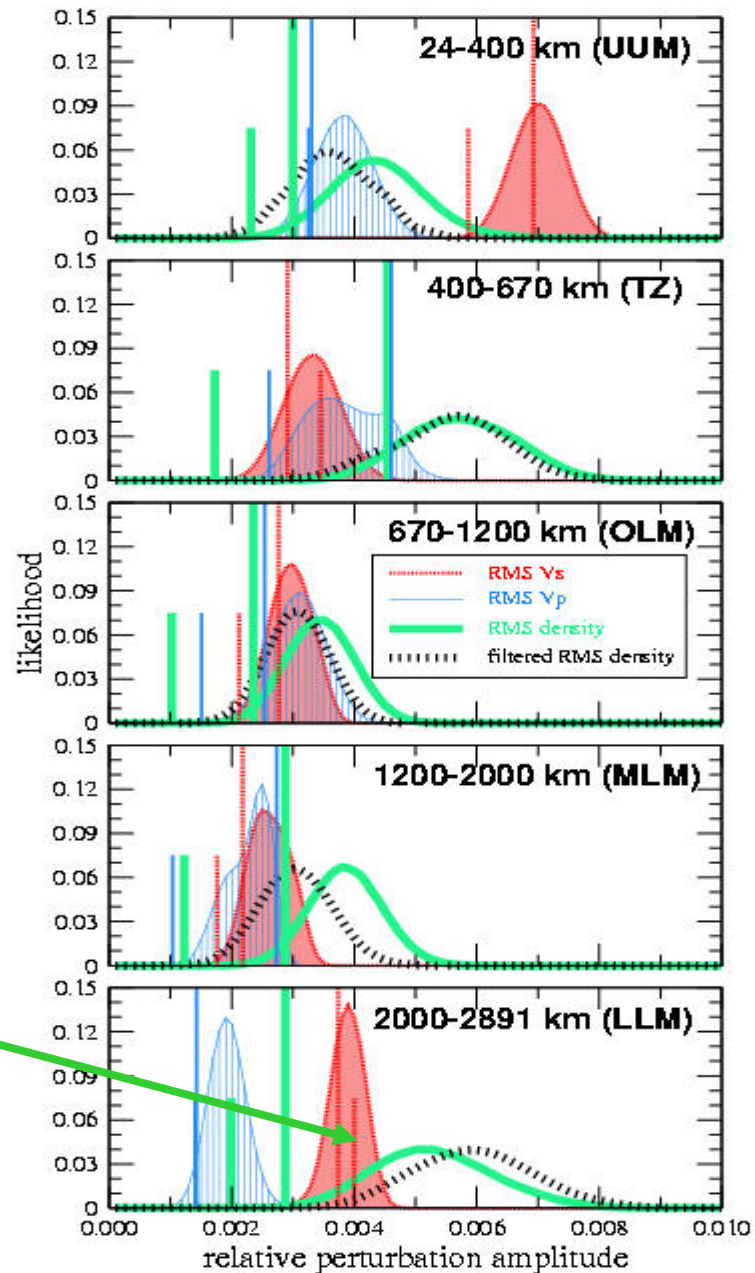
Likelihood of correlation between seismic parameters

Evidence for
Half-height vertical bar
chemical heterogeneity
corresponds to SPRD6
 $V_s - V_\Phi$



Likelihood of rms-amplitude of seismic parameters

Evidence for chemical heterogeneity
RMS density



The neural network (NN) approach:

Bishop 1995, MacKay 2003

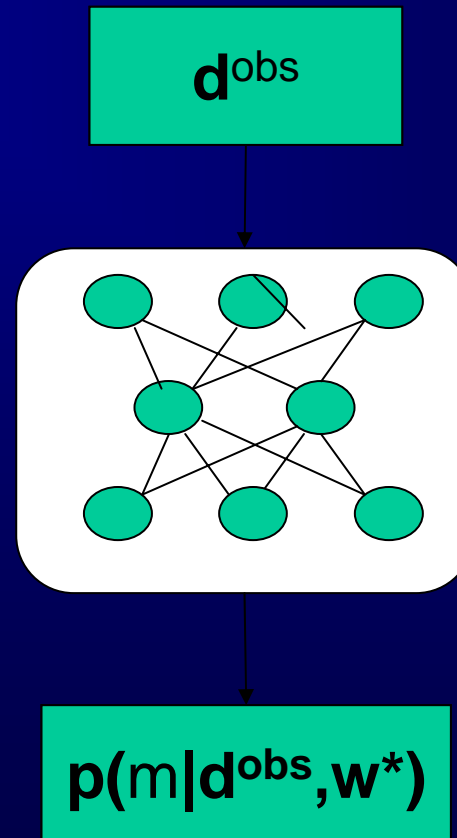
- A neural network can be seen as a non-linear filter between some input and output
- The NN is an approximation to any function g in the non-linear relation $d=g(m)$
- A training set is used to calculate the coefficients of the NN by non-linear optimisation

3) Forward propagating a new datum through the trained network (i.e. solving the inverse problem)

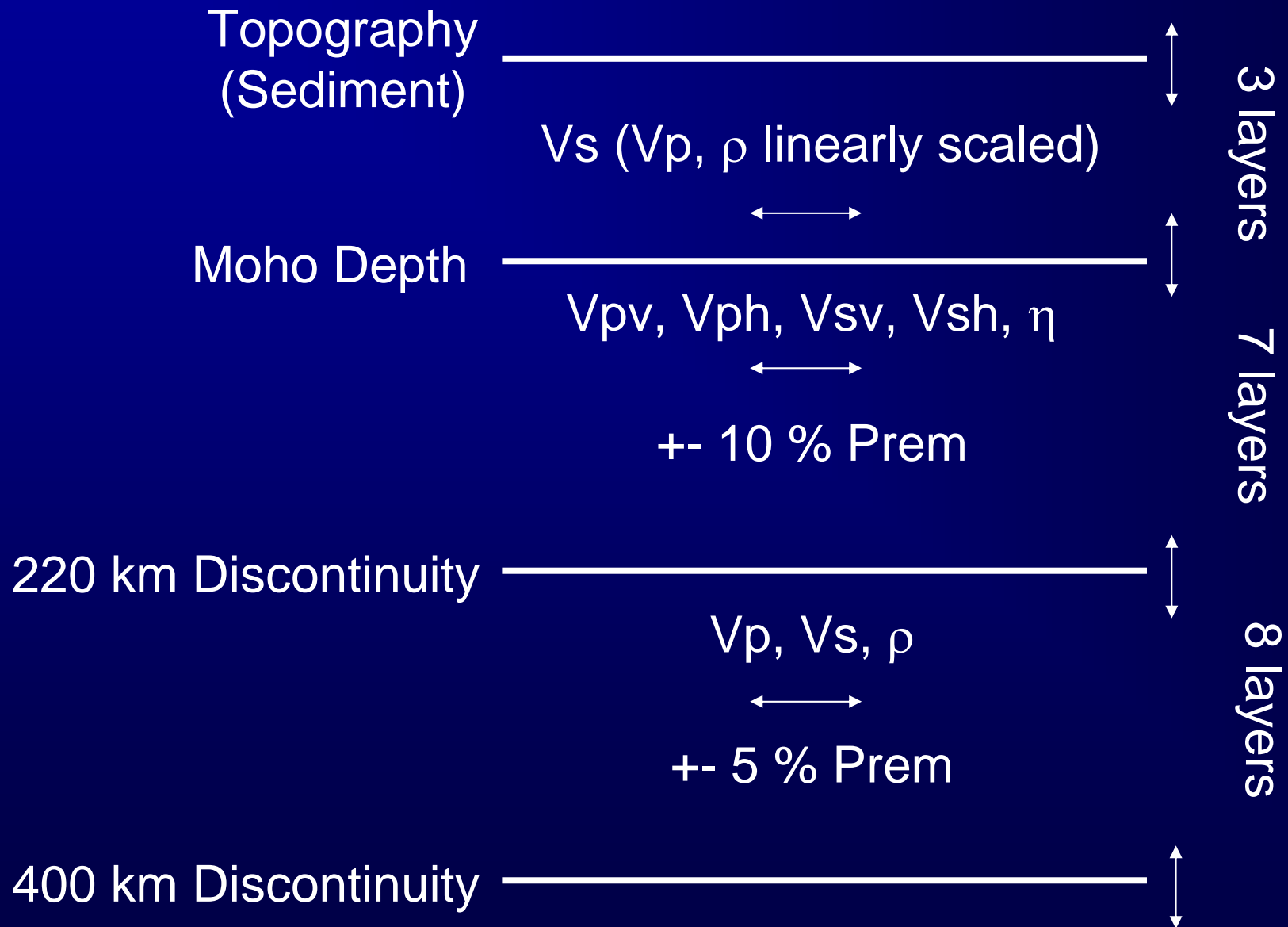
Network Input:
Observed or synthetic
data for training

Neural Network

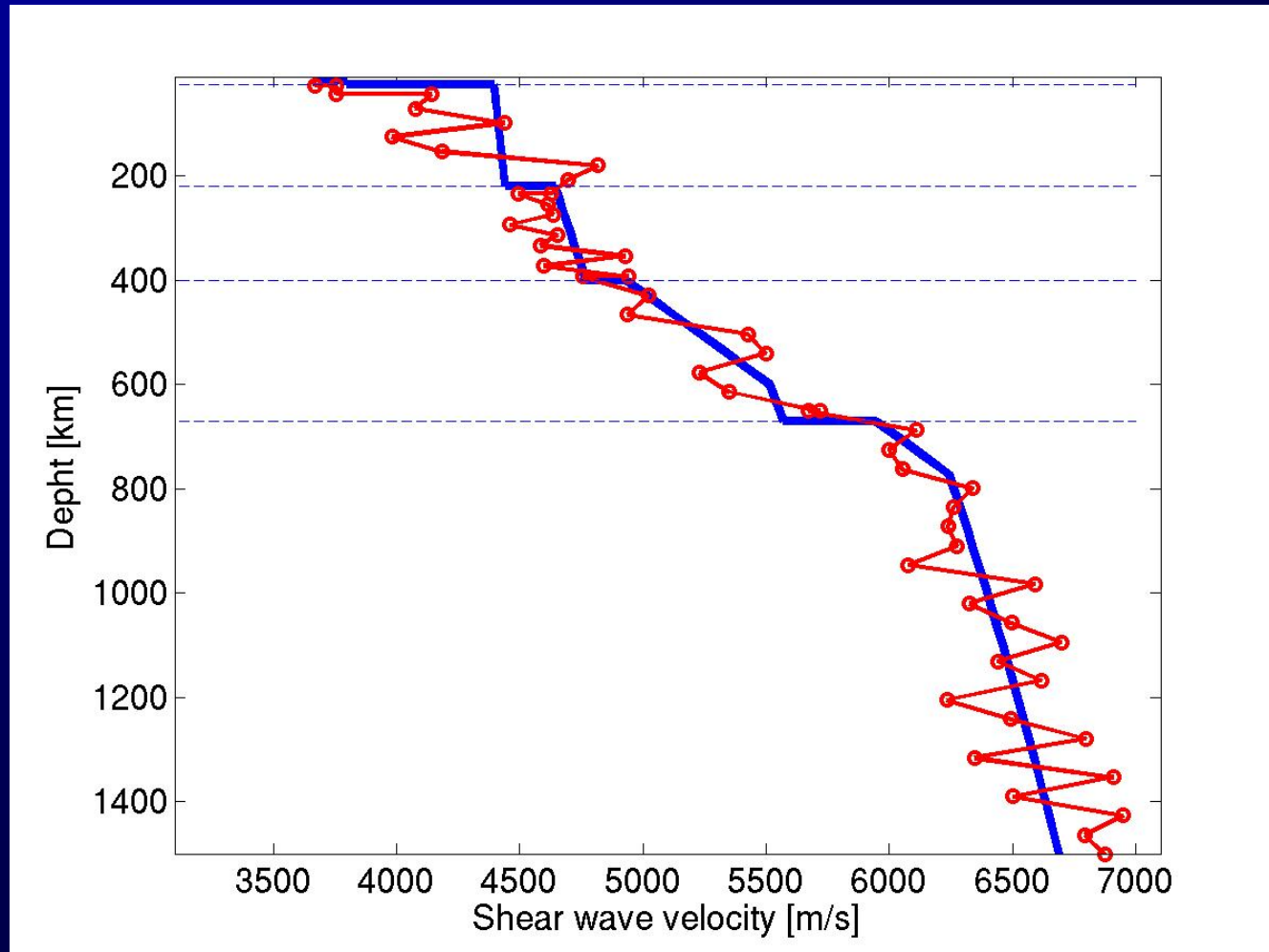
Network Output:
Conditional probability
density



Model Parameterization

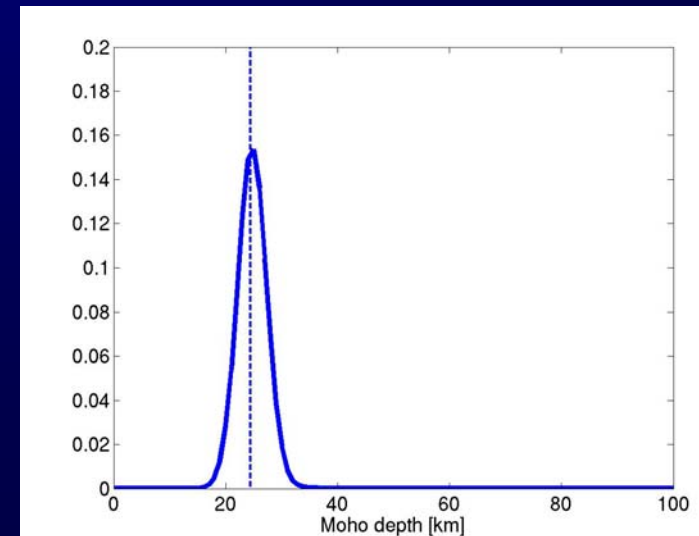


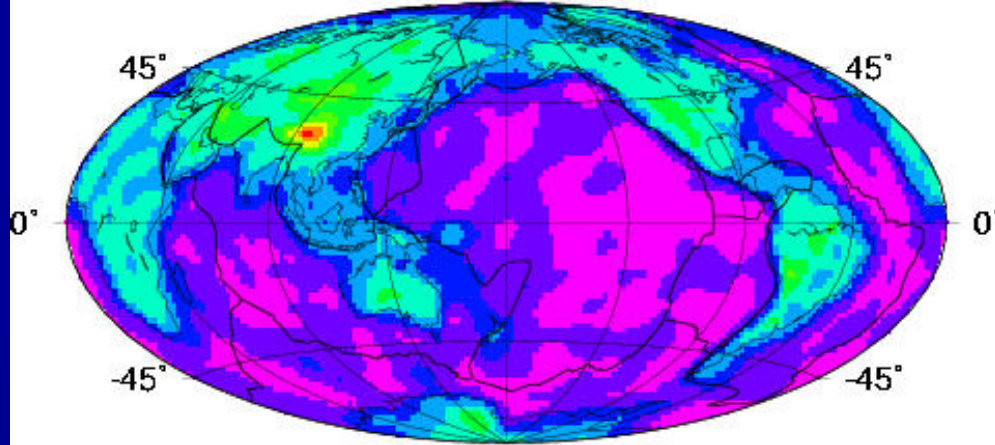
Example of a V_s realisation



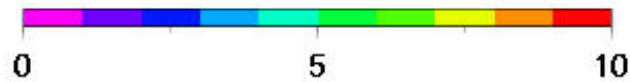
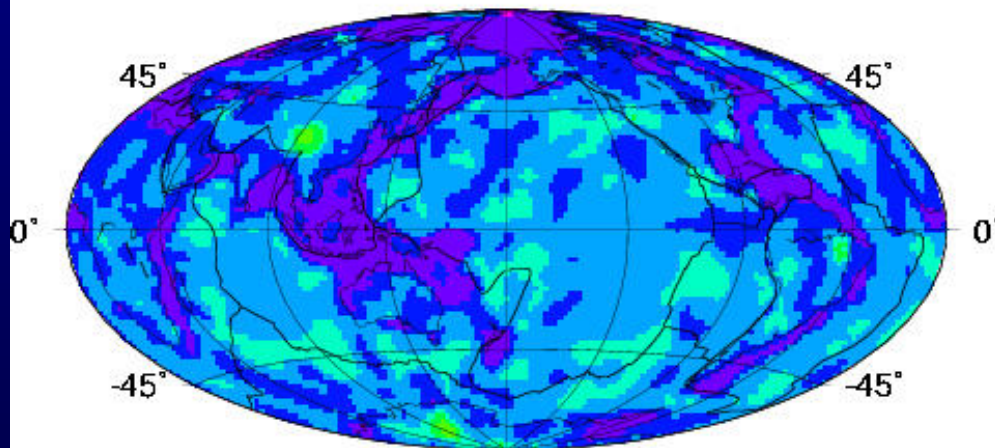
Advantages of NN

1. The curse of dimensionality is not a problem because NN approximates a function not a data prediction!
2. Flexible: invert for any combination of parameters
3. (Single marginal output)



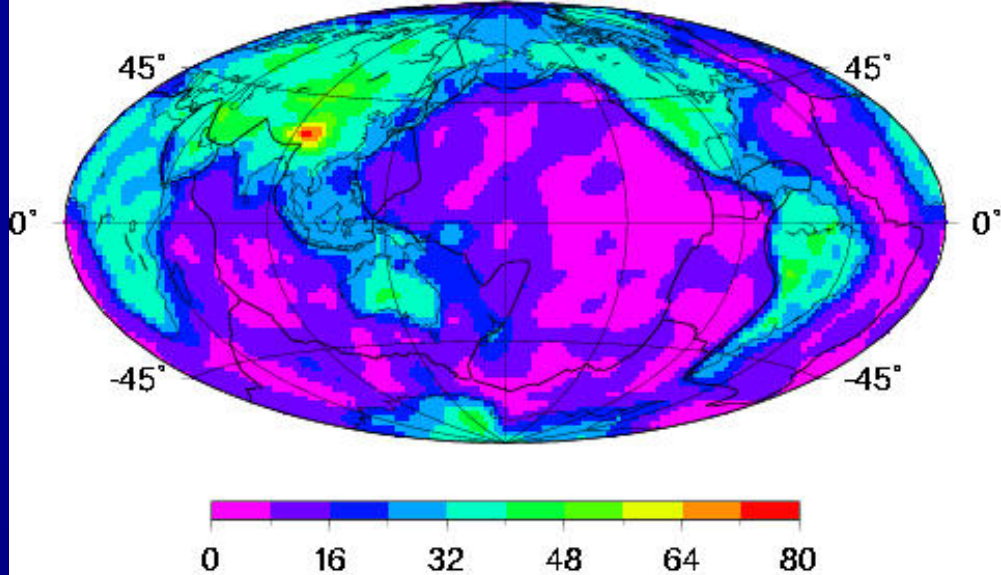


Mean Moho depth [km]

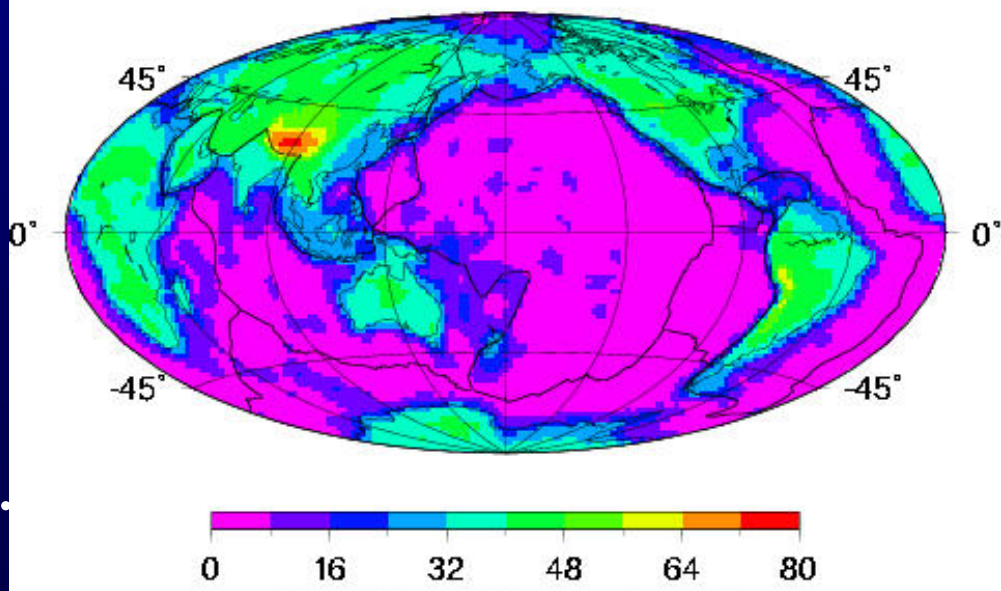


Standard deviation σ [km]

Meier et al.
2007



Mean Moho depth [km]

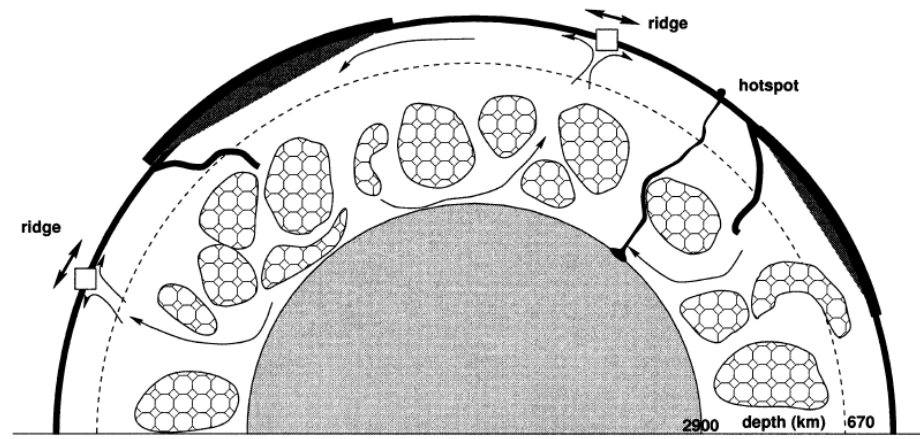
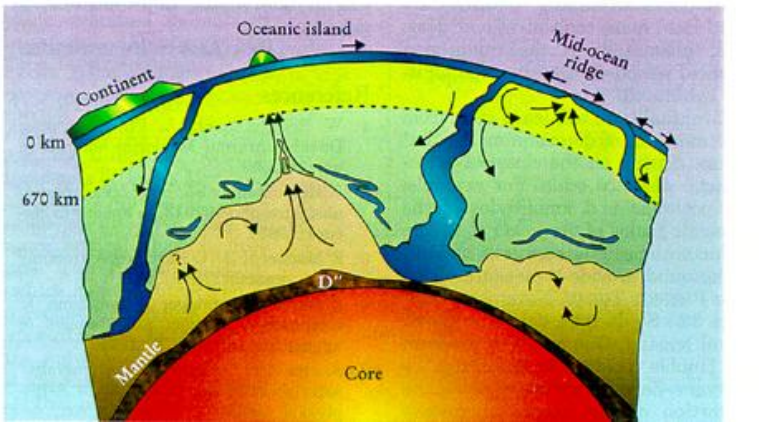


CRUST2.0 Moho depth [km]

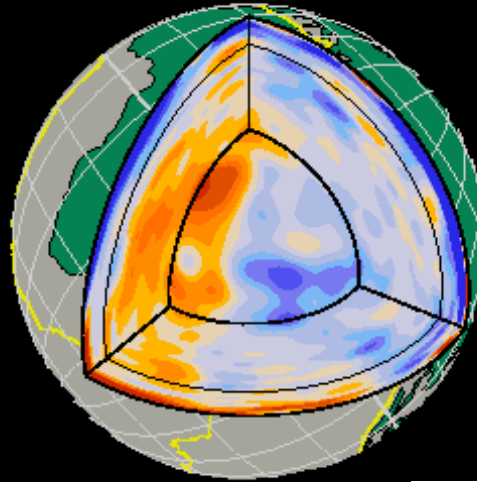
Meier et al.
2007

NA versus NN

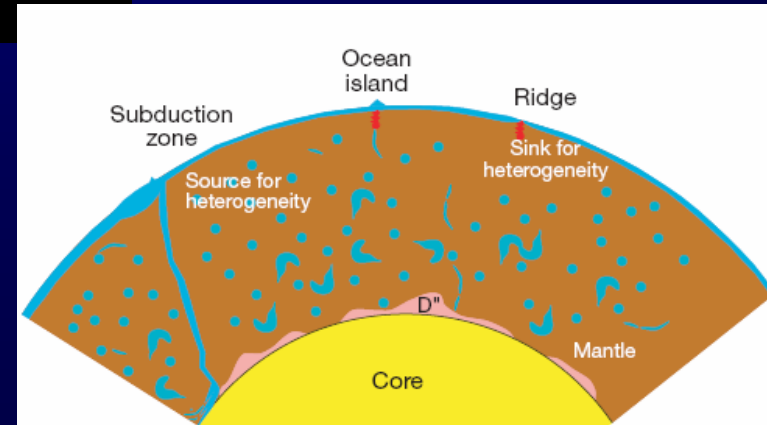
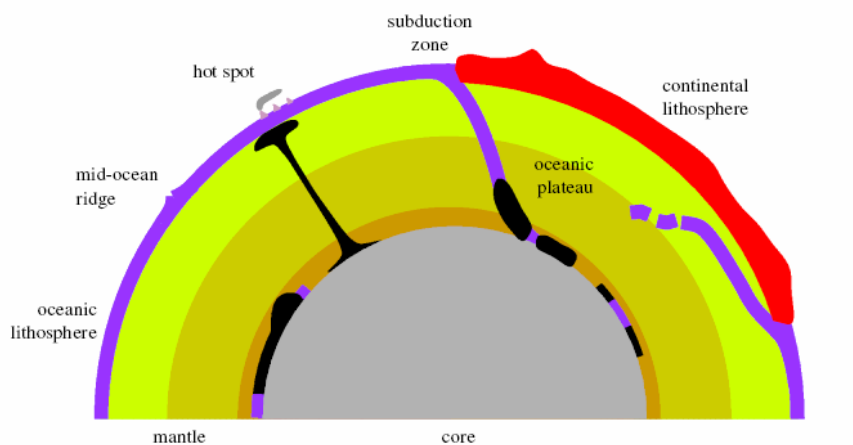
- NA: joint pdf \rightarrow marginals by integration
 \rightarrow small problem
- NN: 1D marginals only (one network per parameter) \rightarrow large problem



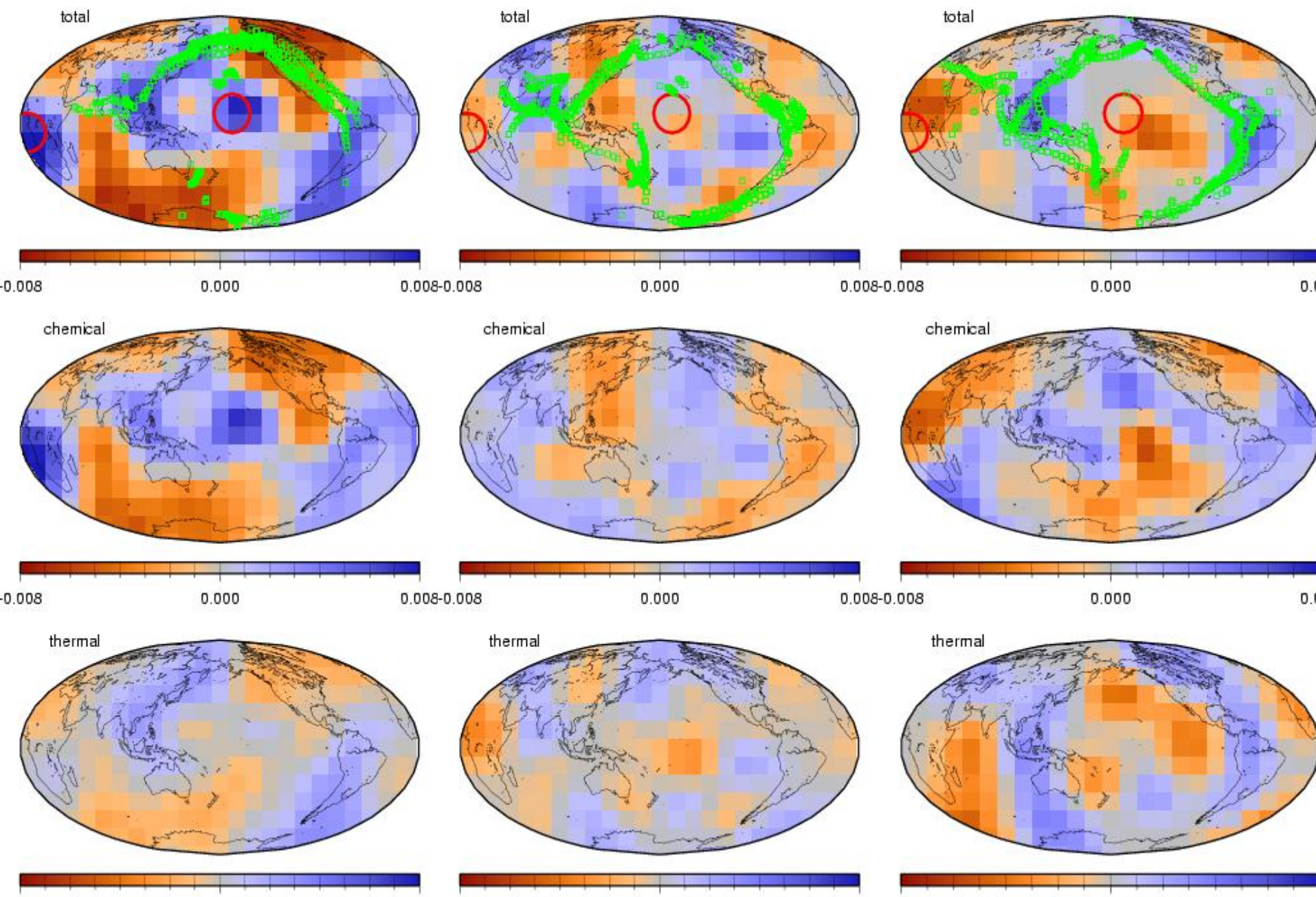
**Advantage
of having a pdf**



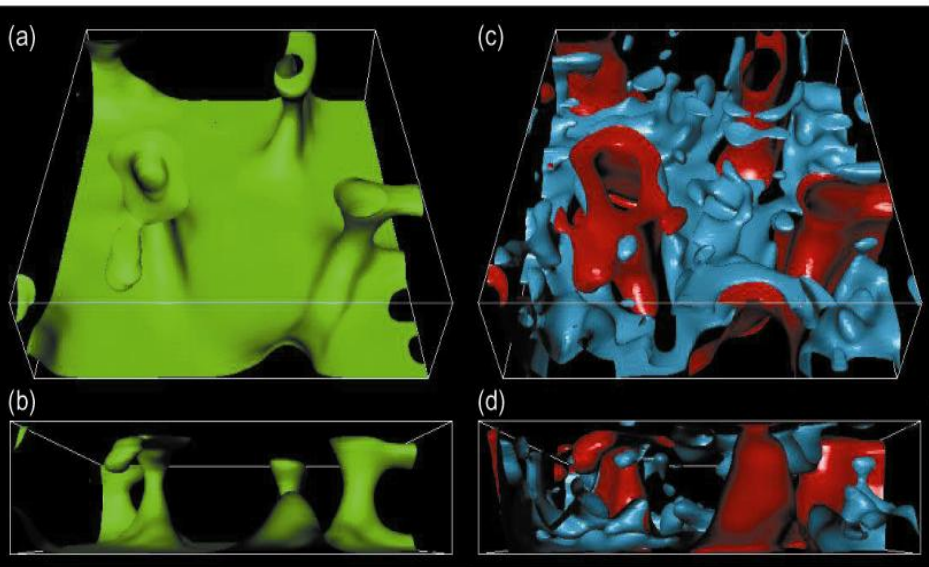
**Quantitative
comparison**



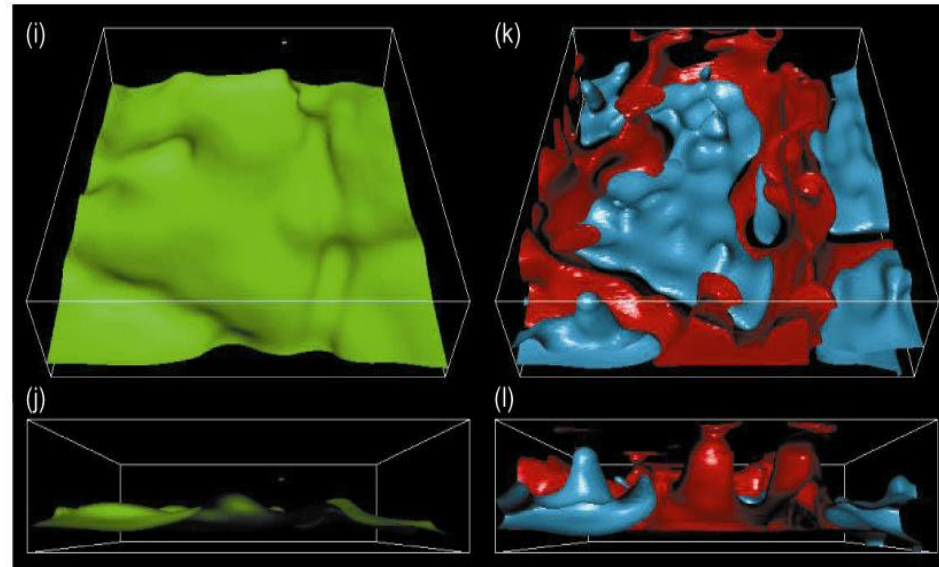
Density 2000-2891 km, gravity filtered Density 1200-2000 km, gravity filtered Density 670-1200 km, gravity filtered



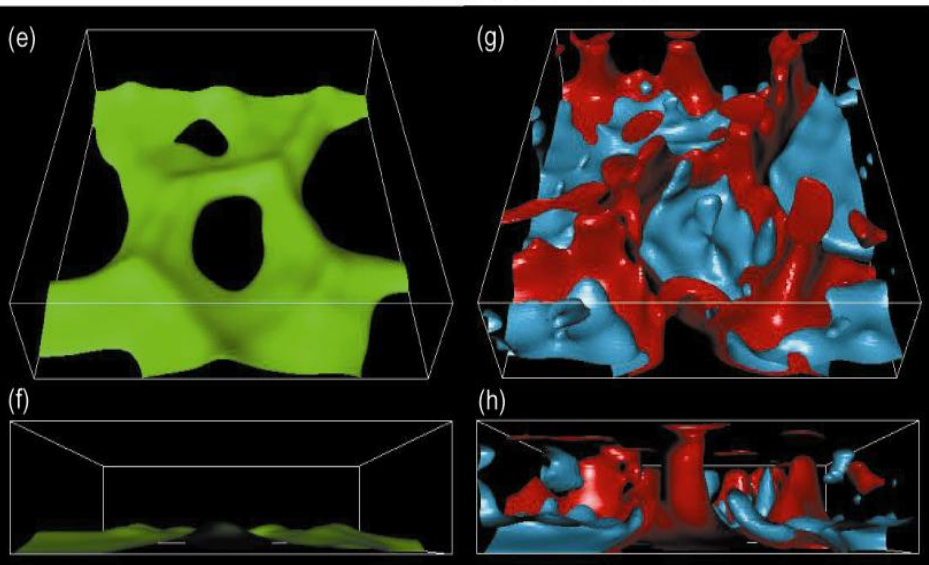
$B=0.2; X=10\%$



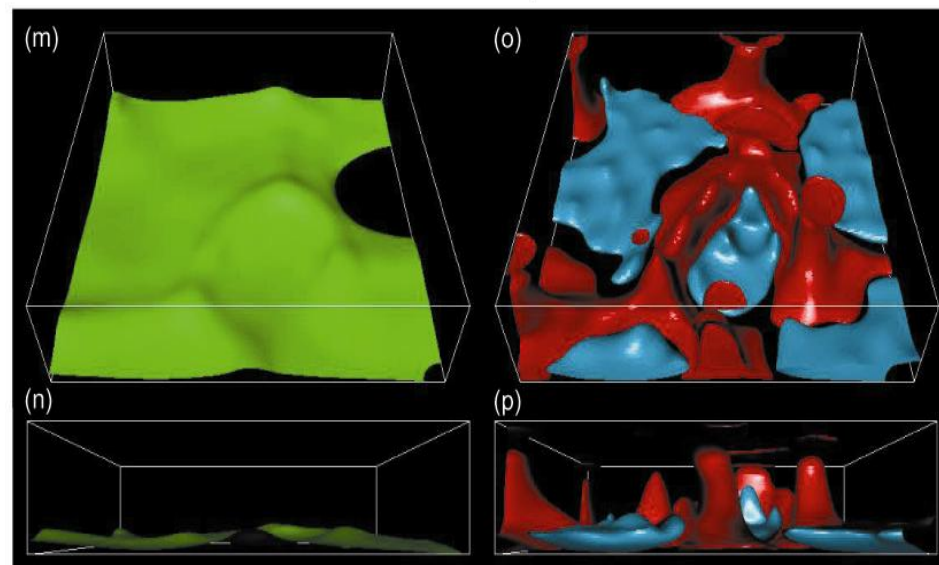
$B=0.3; X=30\%$

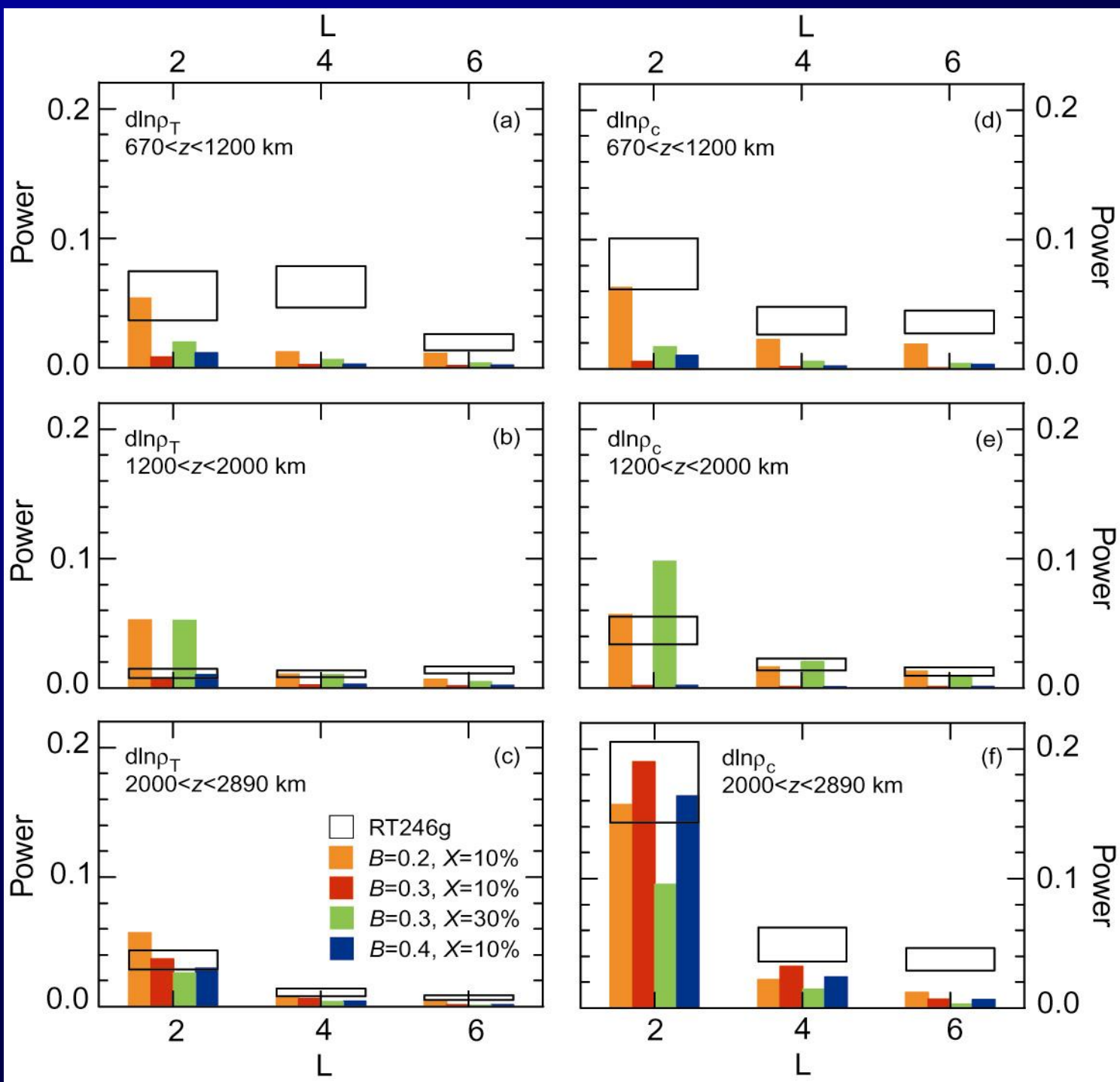


$B=0.3; X=10\%$



$B=0.4; X=10\%$





Finally let's add some SPICE

- For years to come, full Monte Carlo together with 3D wave propagation is out of reach
- NA: iterative linearized inversion using adjoint kernels (Sieminski, Liu, Tromp) → NA around model of last iteration
- NN: data base of all 3D simulations done worldwide can be used for training.